

Radiative transfer model (VIS – TIR) for natural and urban surfaces

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** DART simulated Mont-Blanc sunrise image*



Outline

1. Introduction
2. Physical principles
3. Surface 3D RTMs
4. Applications
5. Perspectives



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Outline

1. Introduction

2. Physical principles

3. Surface 3D RTMs

4. Applications

5. Perspectives

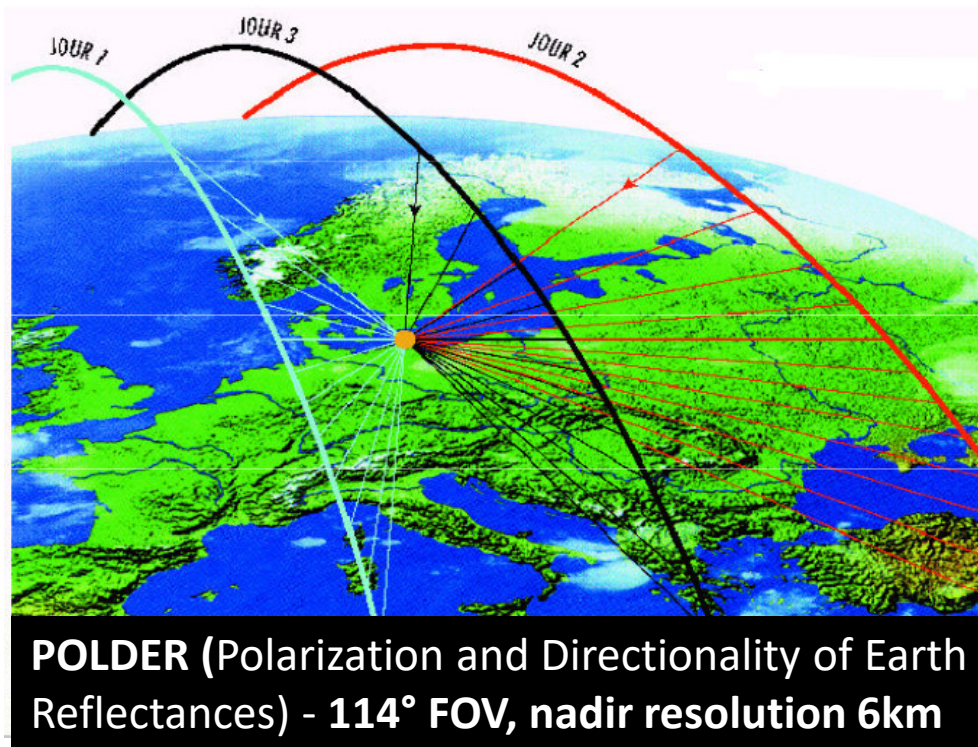


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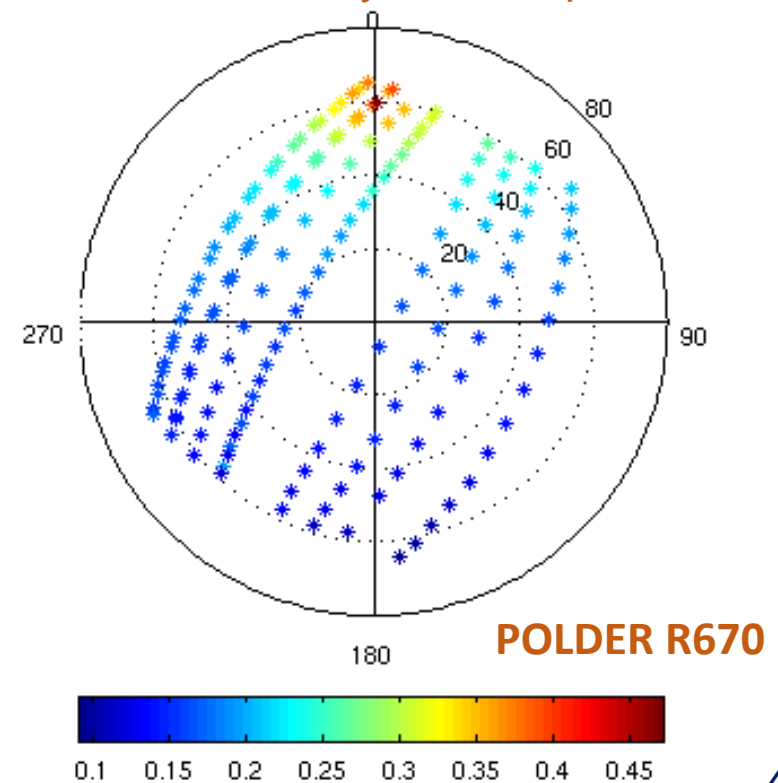
Remote Sensing & Challenges

Directional effect: Observed reflectance ρ_{BOA} changes with Ω_s, Ω_v



*(E. Vermote, C. Justice and Breon, NASA
supported Land LTDR Project)*

incident direction Ω_s : local time
view direction Ω_v : sensor position

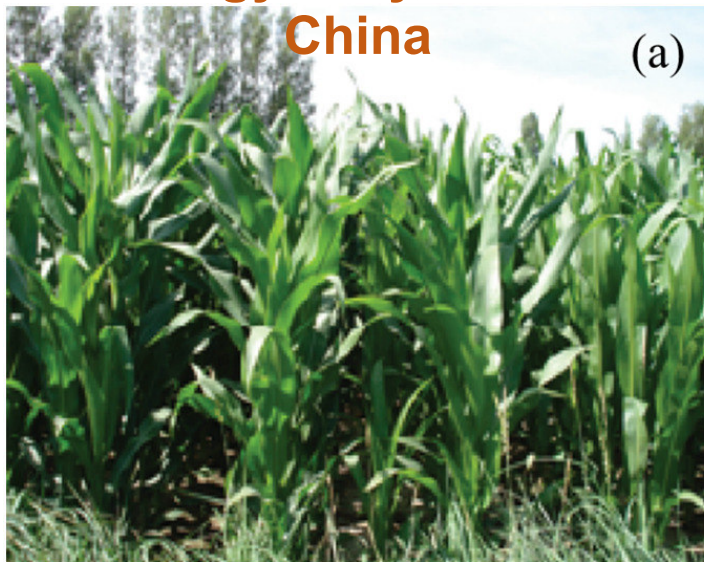




Remote Sensing & Challenges

Directional effect: Observed thermal radiance BT_{BOA} changes with Ω_s, Ω_v

**Zhangye city, Gansu,
China**

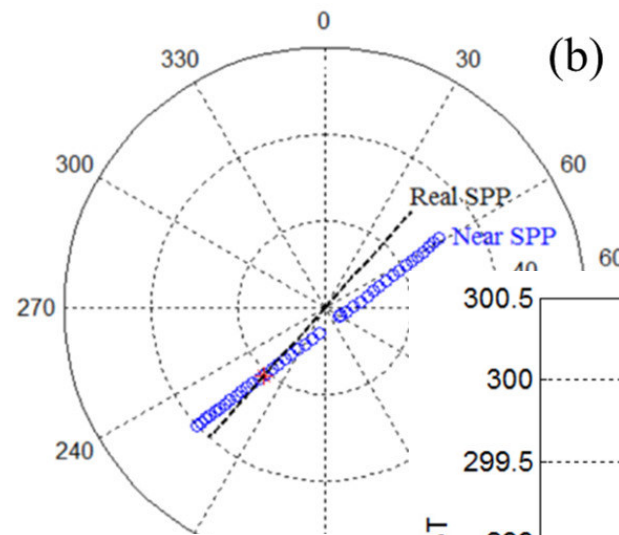


(a)

Airborne TIR camera

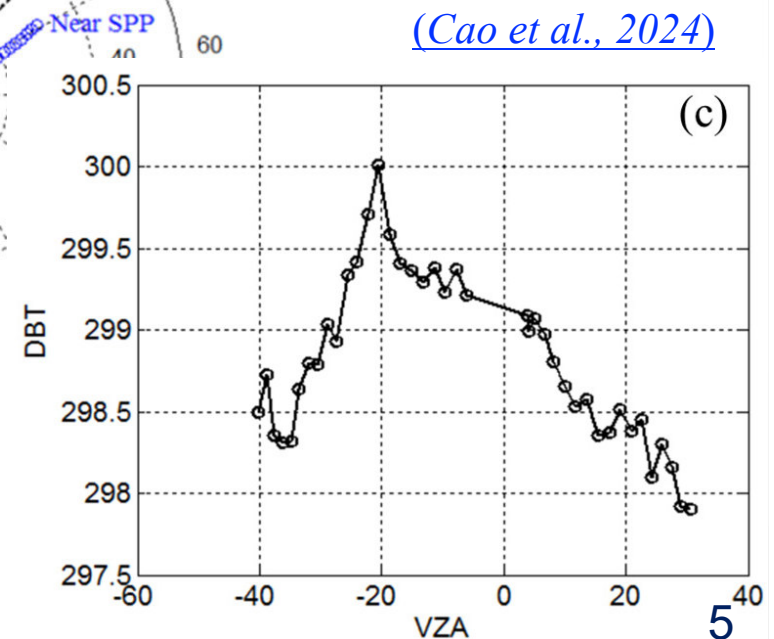
40° wide FOV

Spectral domain 8-11 μm



(b)

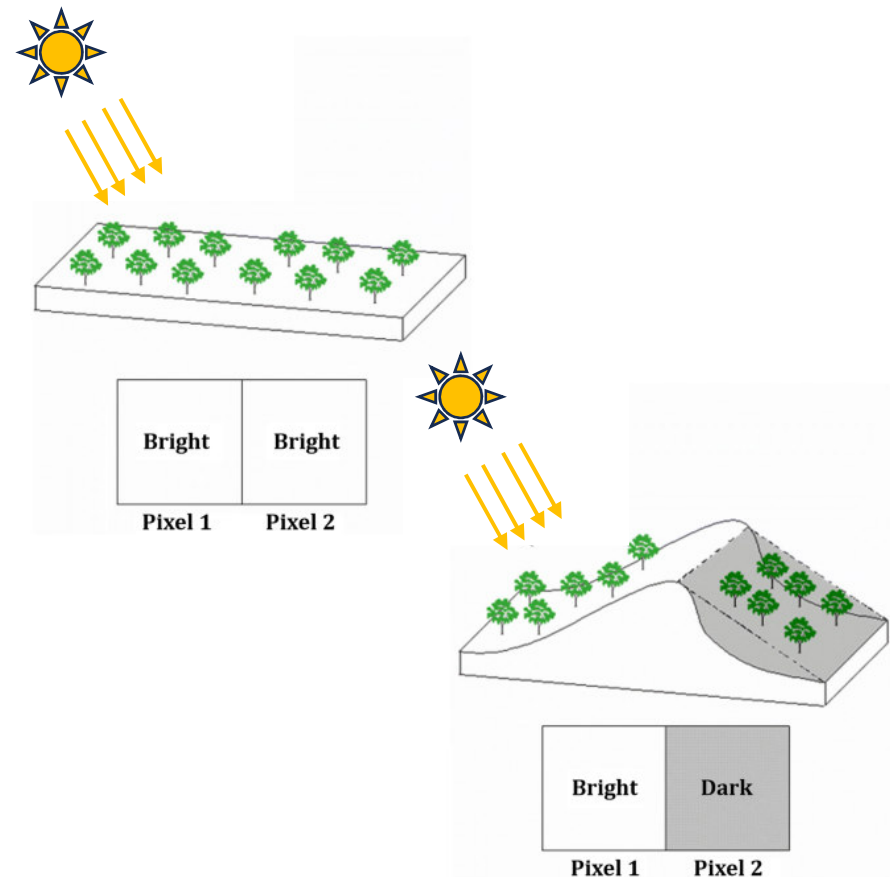
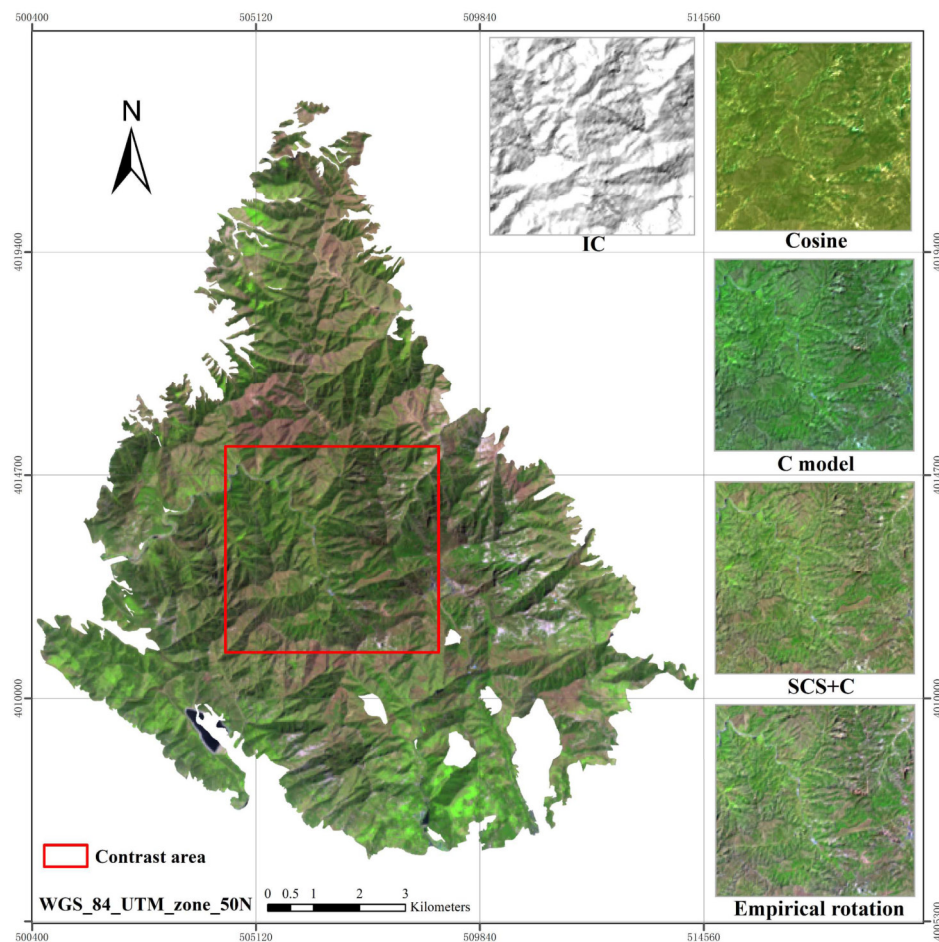
**Near Solar Plane
observation**





Remote Sensing & Challenges

Topographic effect: Observation changes with slope aspect

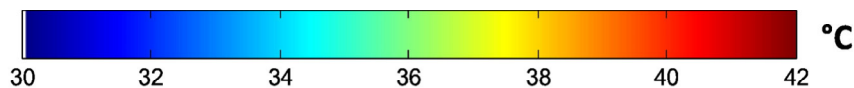
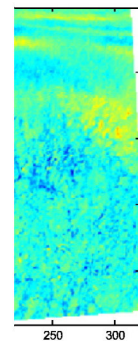
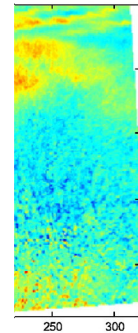
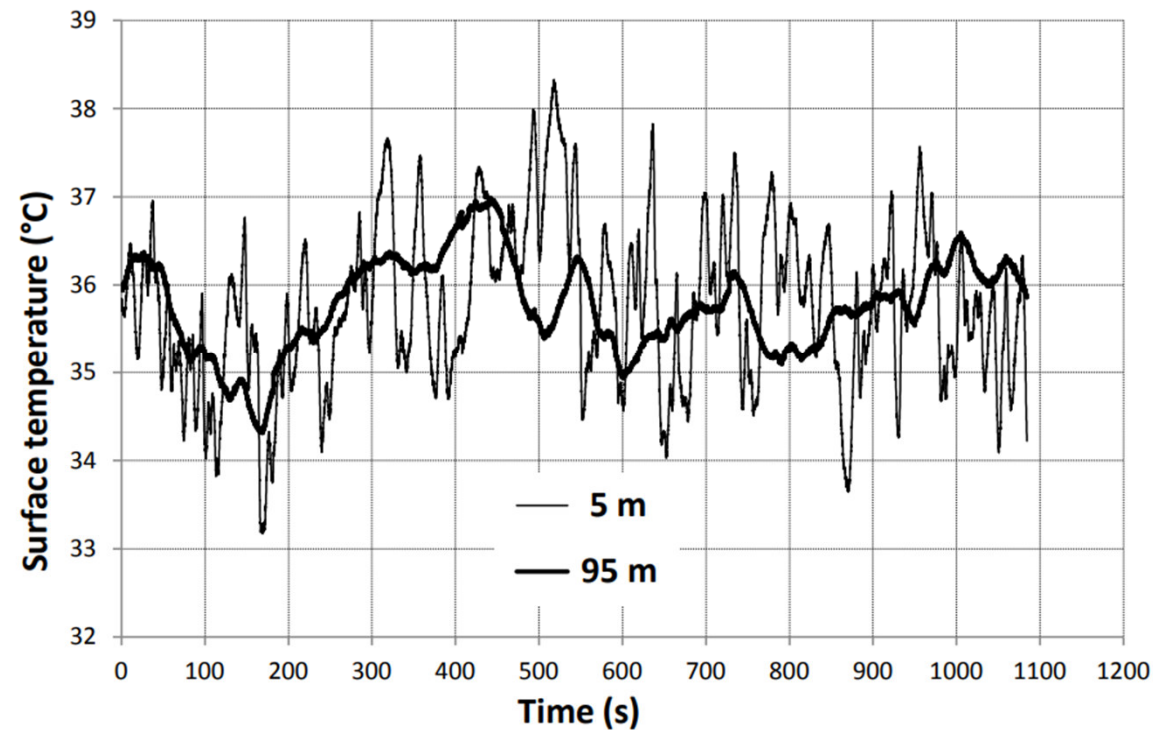
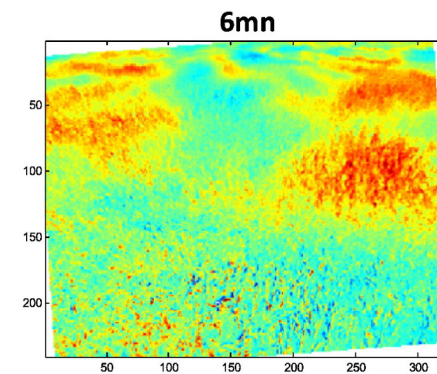


(Chao et al., 2020)



Remote Sensing & Challenges

Natural processes: Observation changes with local microclimate

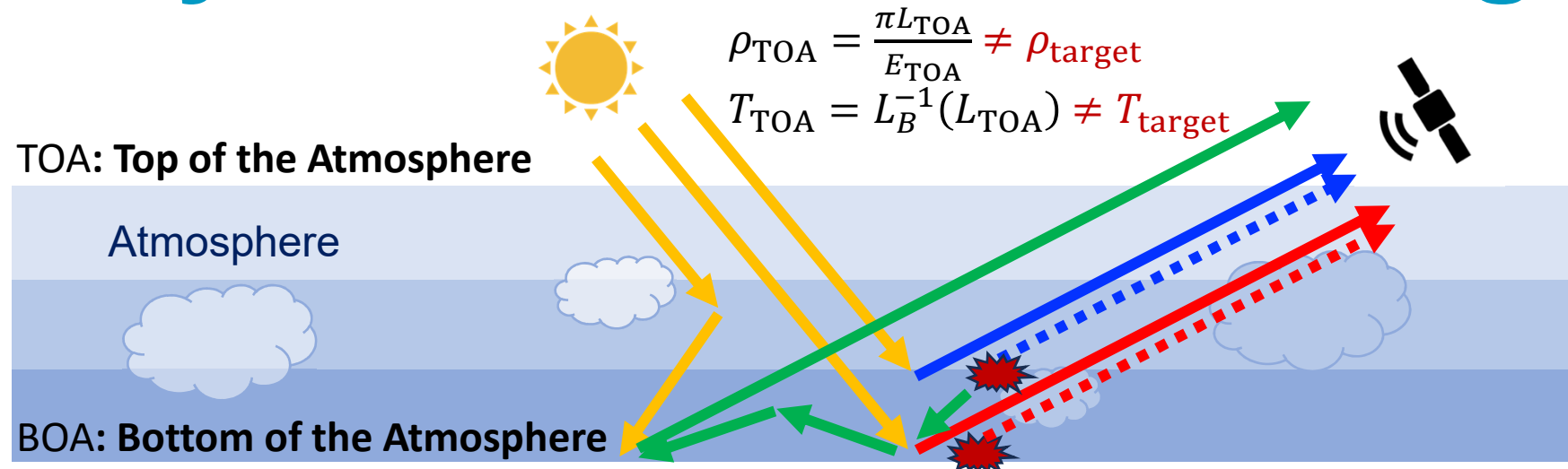


→ Turbulence induced errors on
LST

(Lagouarde et al., 2015)



Physics of Remote Sensing



L_{TOA} = Radiance due to scattering of solar flux and emission by the Earth (Only)
 + Radiance due to scattering of solar flux and emission by the Atmosphere (Only)
 + Radiance due to scattering of flux by {Earth + Atmosphere} (Multiple scattering)

Physical models

Forward: $\text{RS} = f(\pi_1, \pi_2, \dots)$

Inverse: $\text{VAR } \pi_i = f^{-1}(\text{RS})$



Design space missions according to scientific objectives

Derive surface variables from remote sensing observation

Study surface RB to understand land processes and climate

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4. Applications

5. Perspectives



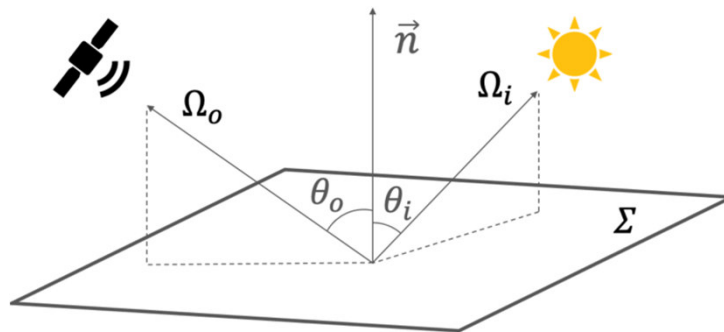
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Physical principles

Light transport equation (surface)

$$L(r, \Omega_o) = L_e(r, \Omega_o) + \int L(r, -\Omega_i) \cdot f(r, \Omega_i, \Omega_o) \cdot |\cos \theta_i| d\Omega_i$$



$$f(\Omega_i, \Omega_o) = \frac{dL(\Omega_i, \Omega_o)}{dE_i(\Omega_i)}$$

$$\int f(r, \Omega_i, \Omega_o) \cdot |\cos \theta_i| d\Omega_i \leq 1$$

$L_e(r, \Omega_o)$: emitted radiance (e.g., light source, thermal emission)

$f(r, \Omega_i, \Omega_o)$: bidirectional scattering distribution function (BSDF)

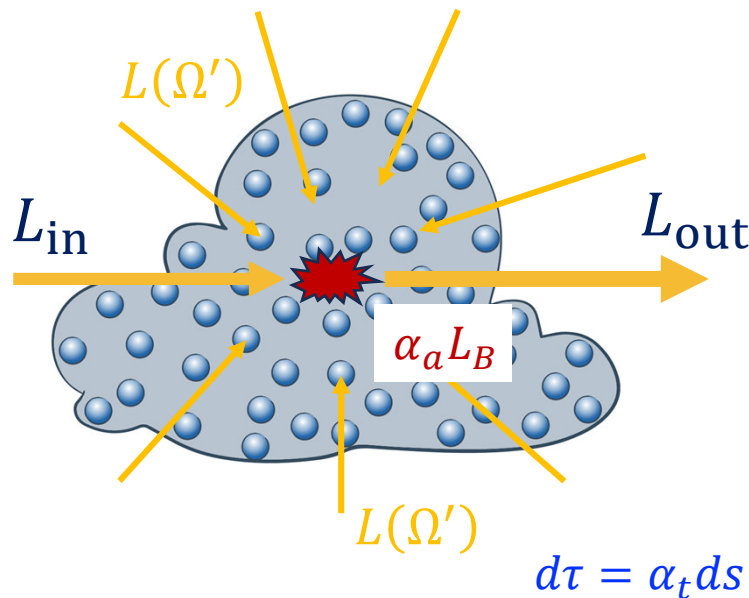
$$f(r, \Omega_i, \Omega_o) = \begin{cases} BRDF: \frac{\rho(r, \Omega_i, \Omega_o)}{\pi}, & \text{if } (\vec{n} \cdot \Omega_i) \cdot (\vec{n} \cdot \Omega_o) \geq 0 \\ BTDF: \frac{\tau(r, \Omega_i, \Omega_o)}{\pi}, & \text{otherwise} \end{cases}$$



Physical principles

Radiative transfer equation RTE (volume: atmosphere, turbid, fluid)

$$\frac{dL(\Omega)}{ds} = -\alpha_t L(\Omega) + \frac{\alpha_s}{4\pi} \cdot \int_{4\pi} L(\Omega') \cdot P(\Omega' \rightarrow \Omega) d\Omega' + \alpha_a L_B(T, \Omega)$$



$-\alpha_t L$: Attenuation due to **absorption** and **out-scattering**

$\frac{\alpha_s}{4\pi} \cdot \int_{4\pi} L(\Omega') \cdot P(\Omega' \rightarrow \Omega) d\Omega'$: Contribution due to **in-scattering** into the line of sight

$\alpha_a L_B$: Contribution due to **thermal emission**

Also expressed as:
$$\frac{dL(\Omega)}{d\tau(\Omega)} = -L(\Omega) + \frac{\omega}{4\pi} \int_{4\pi} L(\Omega') P(\Omega' \rightarrow \Omega) d\Omega' + (1 - \omega) L_B(T, \Omega)$$

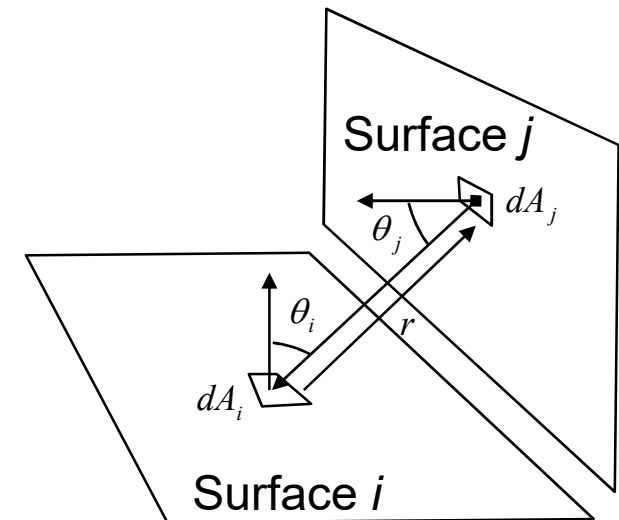


Numerical solution

Radiosity method: Radiant energy conservation.

$$B_i = E_i + \sum_{j=1, j \neq i}^N \xi_i \cdot F_{i,j} \cdot B_j, \text{ with } \xi_i = \rho_i \text{ or } \tau_i$$

$$F_{ij} = \frac{1}{A_i} \int_{A_i} \int_{A_j} \frac{\cos \theta_i \cos \theta_j}{\pi |r|^2} dA_i dA_j$$



B_i : Radiant flux leaving a facet i (N facets in total)

E_i : Emitted flux from a facet i

ρ_i (reflectance): fraction of an incident flux reflected by facet i

τ_i (transmittance): fraction of an incident flux reflected by facet i

$F_{i,j}$ (view factor): fraction of flux leaving facet j and reaching facet i



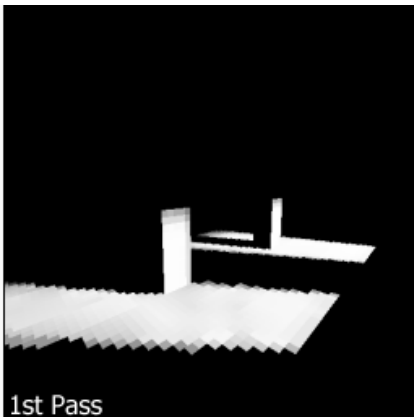
Numerical solution

Radiosity method: Transformation and solution

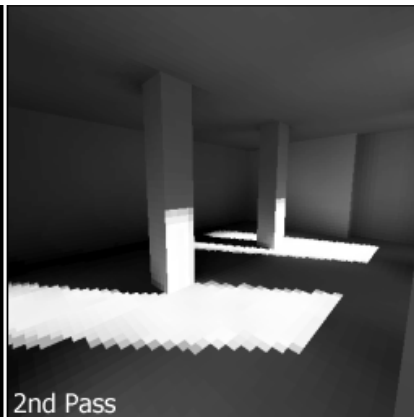
$$B_i = E_i + \xi_i \sum_{j=1, j \neq i}^N F_{i,j} \cdot B_j \quad \rightarrow \quad B_i - \xi_i \sum_{j=1, j \neq i}^N F_{i,j} \cdot B_j = E_i$$

N equations
N unknowns

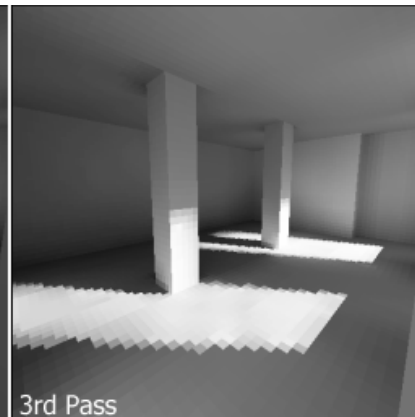
$$\begin{bmatrix} 1 & -\xi_1 F_{1,2} & \dots & -\xi_1 F_{1,N} \\ \vdots & & & \vdots \\ -\xi_N F_{N,1} & -\xi_N F_{N,2} & \dots & 1 \end{bmatrix} \cdot \begin{bmatrix} B_1 \\ \vdots \\ B_N \end{bmatrix} = \begin{bmatrix} E_1 \\ \vdots \\ E_N \end{bmatrix}$$



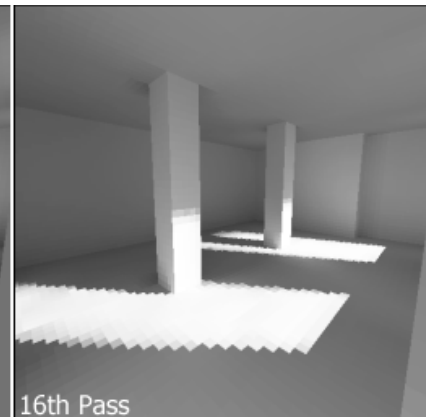
1st Pass



2nd Pass



3rd Pass

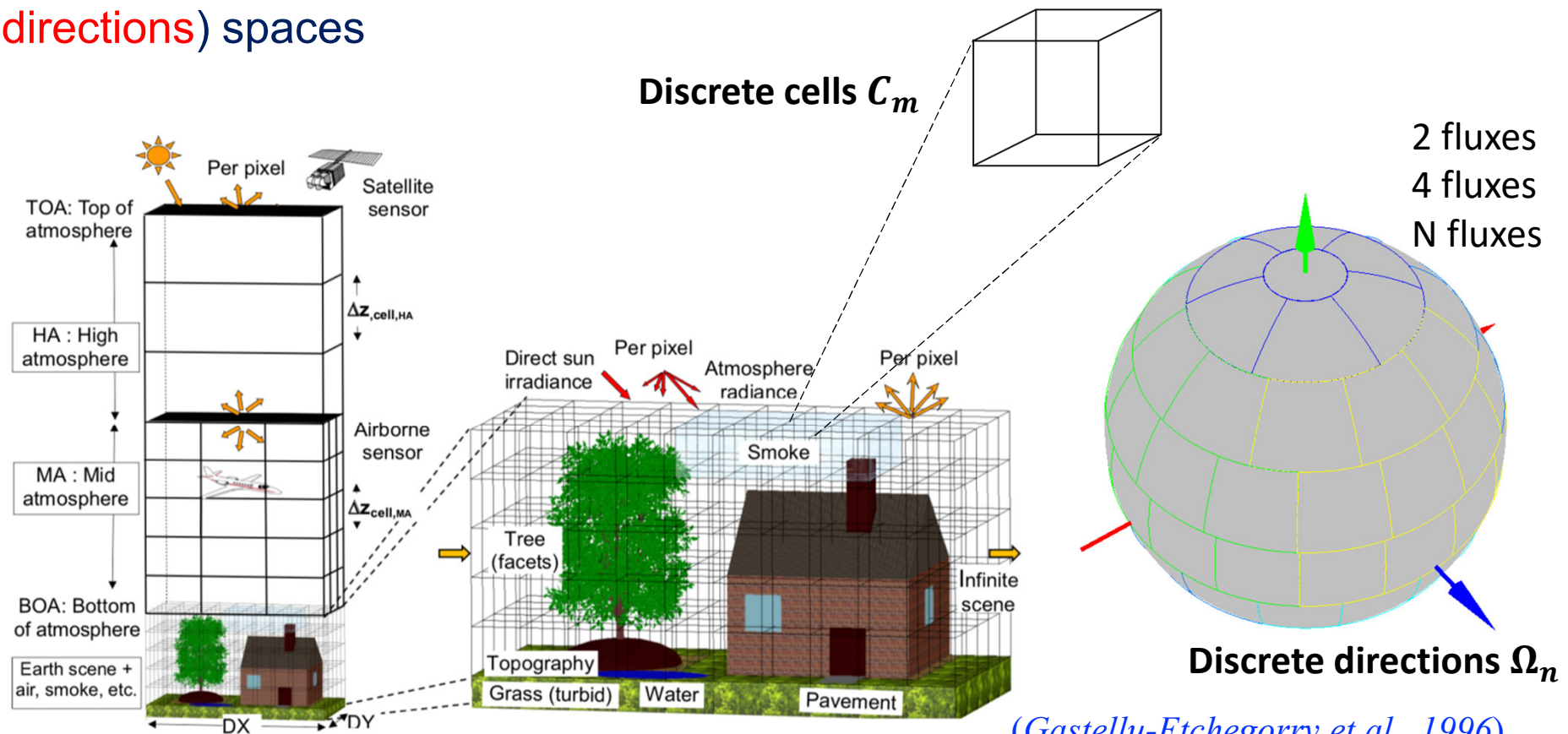


16th Pass



Numerical solution

Discrete Ordiantes method: Discretisation of XYZ (M cells) and angular (N directions) spaces



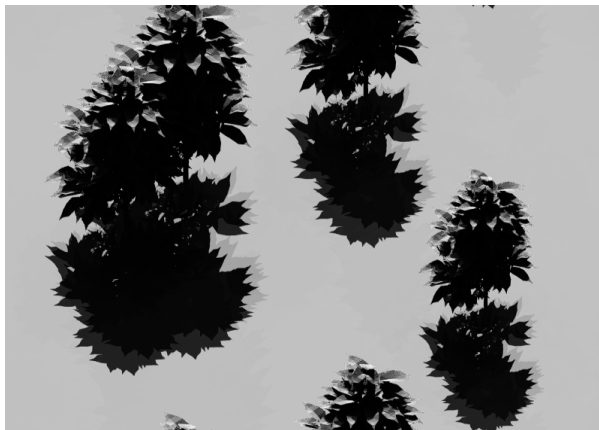


Numerical solution

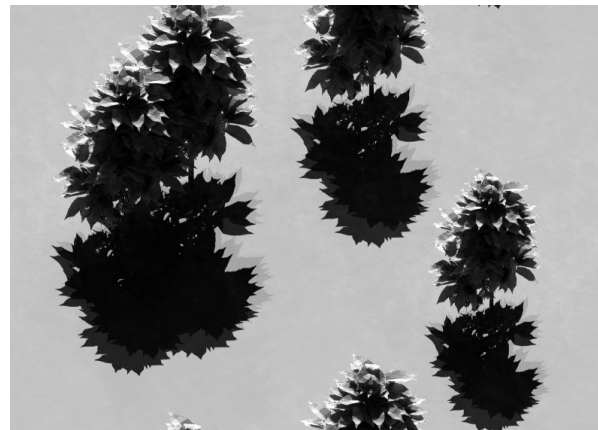
Discrete Ordiantes method: N equations per cell, for M cells

$$L_{\text{out}}(\Omega_m) = \mathcal{T}(\Omega_m)L_{\text{in}}(\Omega_m) + L_e(\Omega_m) + \sum_{i=1}^M \phi_{i,m} L_{\text{in}}(\Omega_i)$$

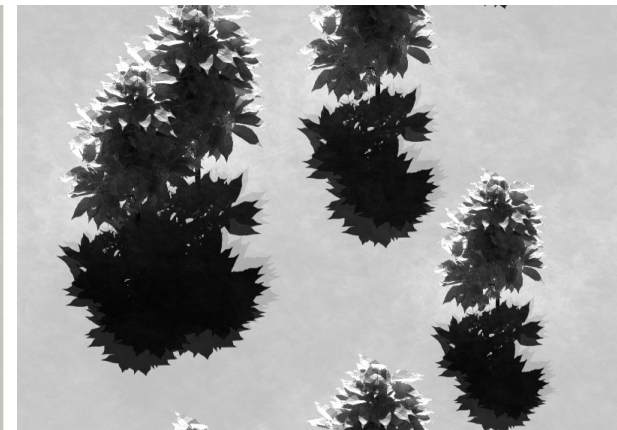
M x N equations per Pass



1st Pass



2nd Pass



20th Pass



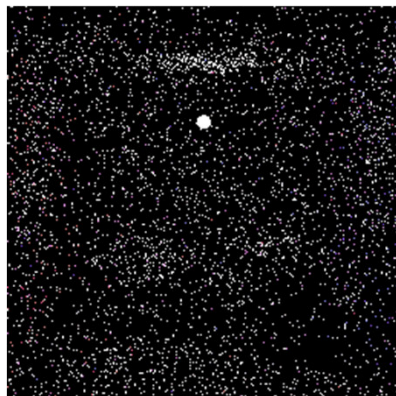
Numerical solution

Monte Carlo method: probability of interception and scattering

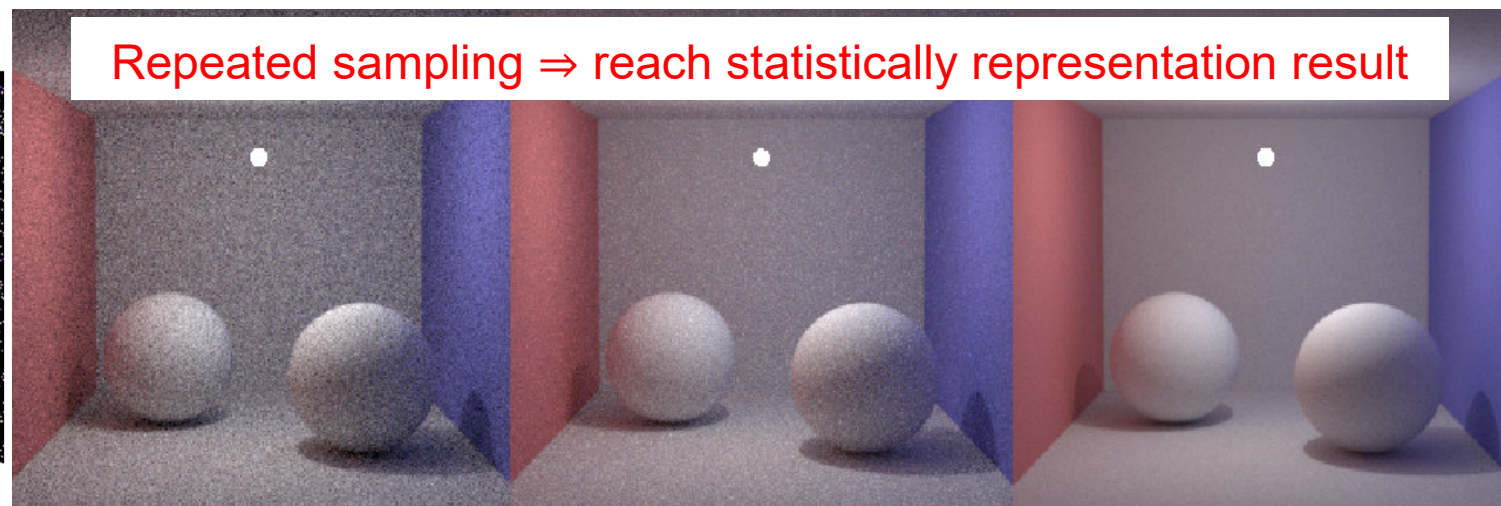
Forward approach: scene illumination by tracking rays from the source

Backward approach: scene illumination by tracking rays from the sensor

Bi-directional approach: scene illumination by tracking rays from sensor & source



MC.gif



Repeated sampling $\Rightarrow$ reach statistically representation result

10 samples / pixel

100 samples / pixel

1000 samples / pixel

Outline

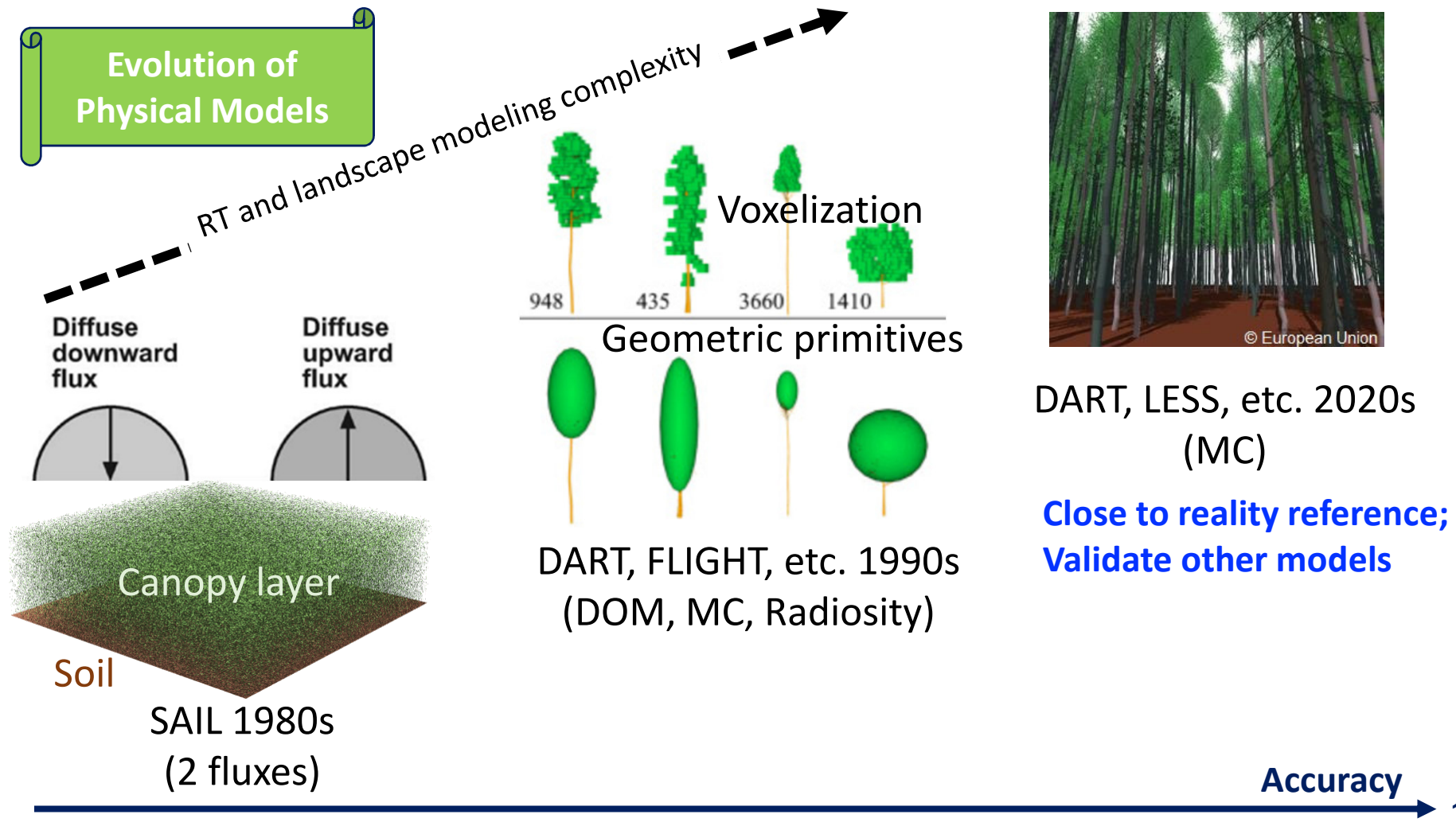
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Radiative transfer modelling



3D radiative transfer models

(Survey on 09/2022, Wang Yingjie, 2022)

■ Yes ■ No ■ Possible

		DART	DIRSIG	Eradiate	FLiES	FLIGHT	LCVRT	LESS	Librat	MCScene	RAPID	Raytran	WPS
Radiative transfer	Mode (MC, DOM, Radiosity)	M/D	M	M	M	M	M	M	M	M	R	M	M
	Short wave [12]	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes
	Thermal emission [7]	Yes	Yes	No	Yes	No	No	Yes	No	Yes	No	No	Yes
	SIF emission [6]	Yes	No	No	Yes	Yes	Yes	No	No	No	No	No	Yes
	LiDAR [7]	Yes	Yes	No	No	Yes	No	Yes	No	Yes	No	No	Yes
	Surface anisotropic scattering [9]	Yes	Yes	No	No	Yes	Yes	No	No	No	Yes	Yes	Yes
	Surface specular reflection [8]	Yes	Yes	No	No	Yes	Yes	No	No	No	Yes	Yes	Yes
	Polarization [3]	Yes	Possible	No	No	No	Yes	No	No	No	No	No	No
	3D radiative budget [8]	Yes	Yes	No	Yes	No	Yes	No	Yes	No	Yes	Yes	Yes
Light source	Parallel sunlight [12]	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes
	Spherical sun [8]	Yes	Yes	No	Yes	Yes	Yes	Yes	No	No	Yes	No	No
	BOA anisotropic atmosphere [10]	Yes	Possible	No	Yes	Yes	Yes	Yes	Yes	Possible	No	Yes	Yes
	Multi light sources [8]	Yes	Yes	No	No	Yes	No	Yes	Yes	No	Yes	Yes	Yes

[*\(Survey on 09/2022, Wang Yingjie, 2022\)*](#)

		DART	DIRSIG	Eradiate	FLiES	FLIGHT	LCVRT	LESS	Librat	MCScene	RAPID	Raytran	WPS
Sensor	Orthographic camera [9]	Yes	Yes	No	Yes	Yes	No	Yes	Yes	Yes	Yes	No	Yes
	Perspective camera [10]	Yes	Yes	Yes	No	Yes	No	Yes	Yes	Yes	Yes	Yes	Yes
	Hemispheric camera [9]	Yes	Yes	No	Yes	Yes	No	Yes	Yes	Yes	Yes	Yes	No
	Pushbroom [8]	Yes	Yes	No	No	Yes	No	Yes	Yes	Yes	Yes	Yes	No
Land surface	Turbid medium (vegetation) [11]	Yes	Yes	No	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes
	Solid surface (buildings, <i>etc.</i>) [11]	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes	Yes
	Fluids (gas, aerosol, <i>etc.</i>) [7]	Yes	Yes	Yes	No	No	No	No	Yes	No	Yes	Yes	Yes
	Topography [9]	Yes	Yes	Yes	Yes	No	Yes	Yes	Yes	Yes	Yes	Yes	No
	Spherical Earth surface [5]	Yes	Yes	Yes	No	No	No	No	Yes	No	No	No	No
Atmos.	Plane parallel [9]	Yes	Yes	Yes	Yes	No	Yes	No	Yes	Yes	No	Yes	Yes
	Spherical [3]	Yes	No	Yes	No	No	No	No	Yes	No	No	No	No
	3D clouds [3]	Yes	Yes	No	No	No	No	No	Yes	No	No	No	No
Energy balance	Photosynthesis, thermal conduction, <i>etc.</i> [5]	Yes	No	No	Yes	Yes	No	No	No	No	No	Yes	Yes

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DART model (CESBIO, FR)

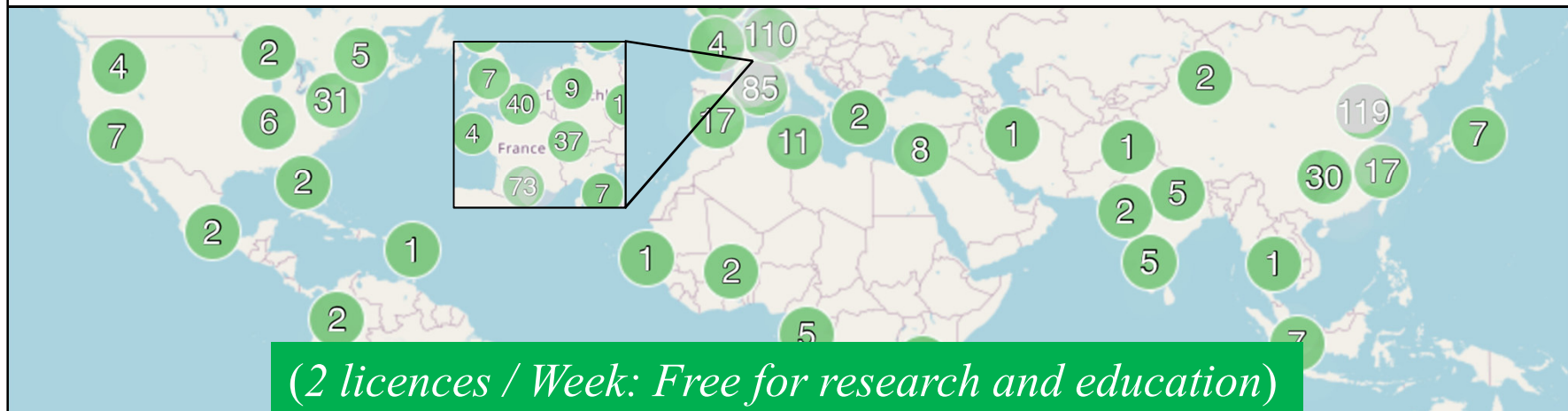
History: developed in CESBIO since 1992 by 10 scientists. Patented in 2003

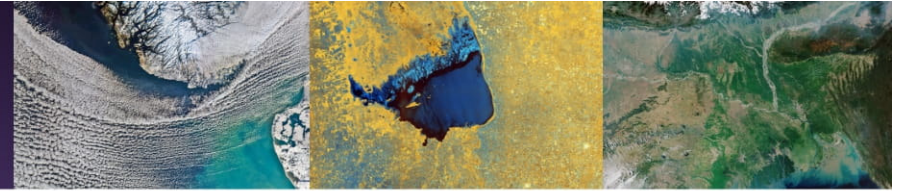
Code: 500,000 lines **C++ (Core)**, 100,000 lines **Java (GUI)** and **Python (Tools)**

Accuracy (relative difference ε , RMSE) assessed with:

- Monte Carlo models (RAMI-III experiment): $\varepsilon_\rho \leq 1\%$ ([Widlowski et al., 2007](#))
- Measurements: $\varepsilon_\rho \leq 2.5\%$ ([Landier et al., 2018](#)), $RMSE_{T_B} < 2K$ ([Sobrino et al., 2011](#))

760 DART licences: Universities, Research centres (CNES, ESA, ...)





Sentinel-2 NG (ESA)

DART scenes used as reference $\Rightarrow$ Trade-off "several satellite" vs "large FOV".
Project Sentinel-2 NG (ESA) *(Gastellu-Etchegorry et al., 2022)*

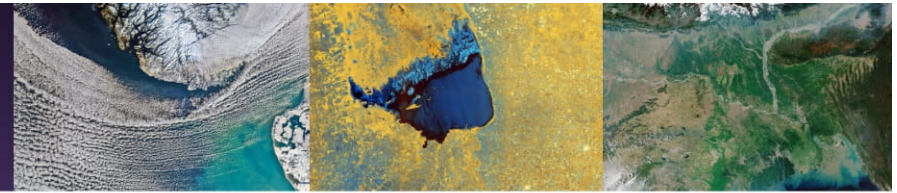


Google map (Ripperdan, USA, Sept. 2018)

3 km



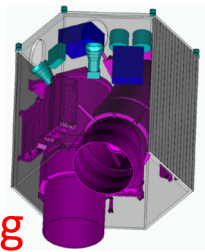
DART Simulation



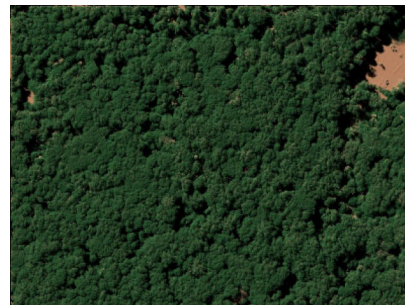
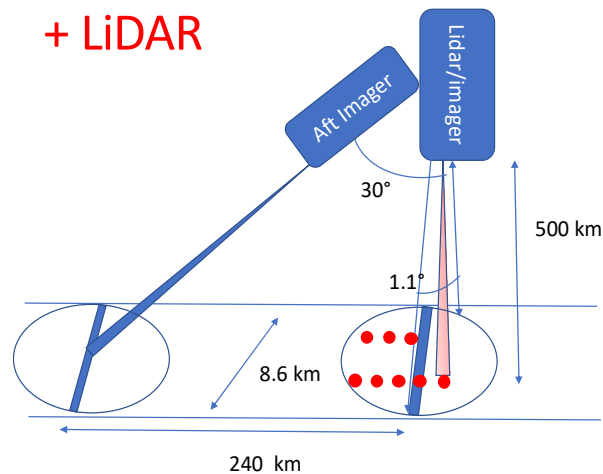
STV (NASA)

Design next generation instrument $\Rightarrow$ HR global topography. Project STV (NASA)

DART gives images and (x, y, z) coordinates of scene elements
 $\Rightarrow$ NASA uses DART to define the optimal satellite configuration



Stereo imaging
+ LiDAR



DART: 400x400m SERC forest



DART: nadir (15/06/2012, 9am)



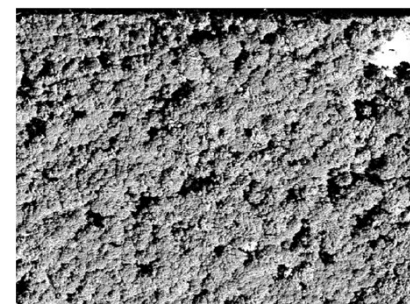
NASA G-LiHT Lidar Reference



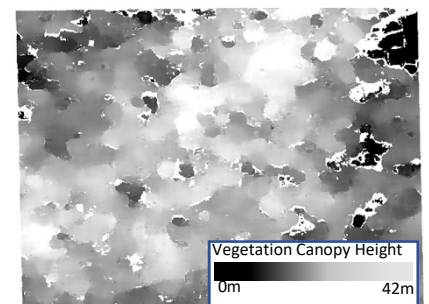
SERC forest



DART: scene-leaf off



DART -20° (15/06/2012, 9am)



Canopy height from DART stereo pair

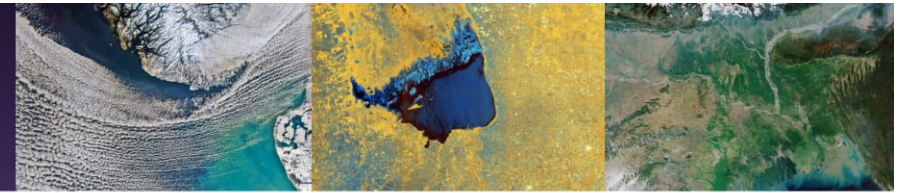


Image processing algorithm

DART Large-Scale Simulations of Stereogrammetry

Google map
(4.5km×1km)



meter-scale

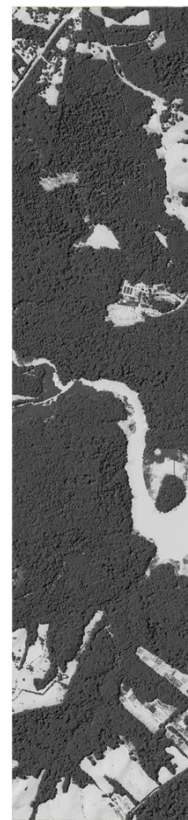
DART simulated images



(a) Nadir



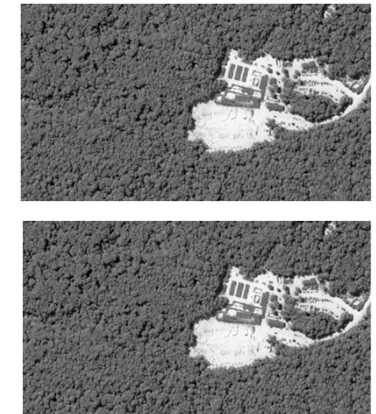
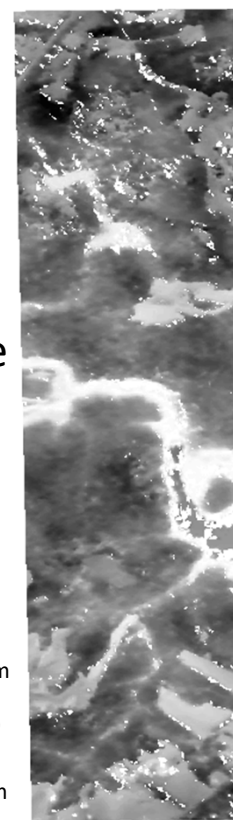
(b) 10° to 40° forward



NASA
ASP
pipeline

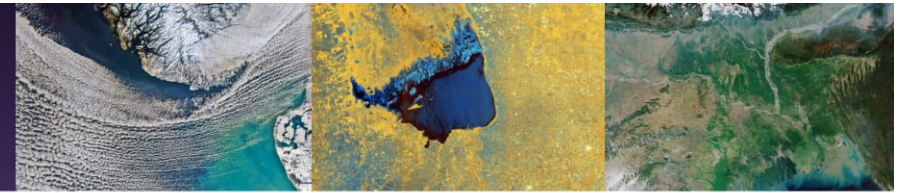


70 m
0 m

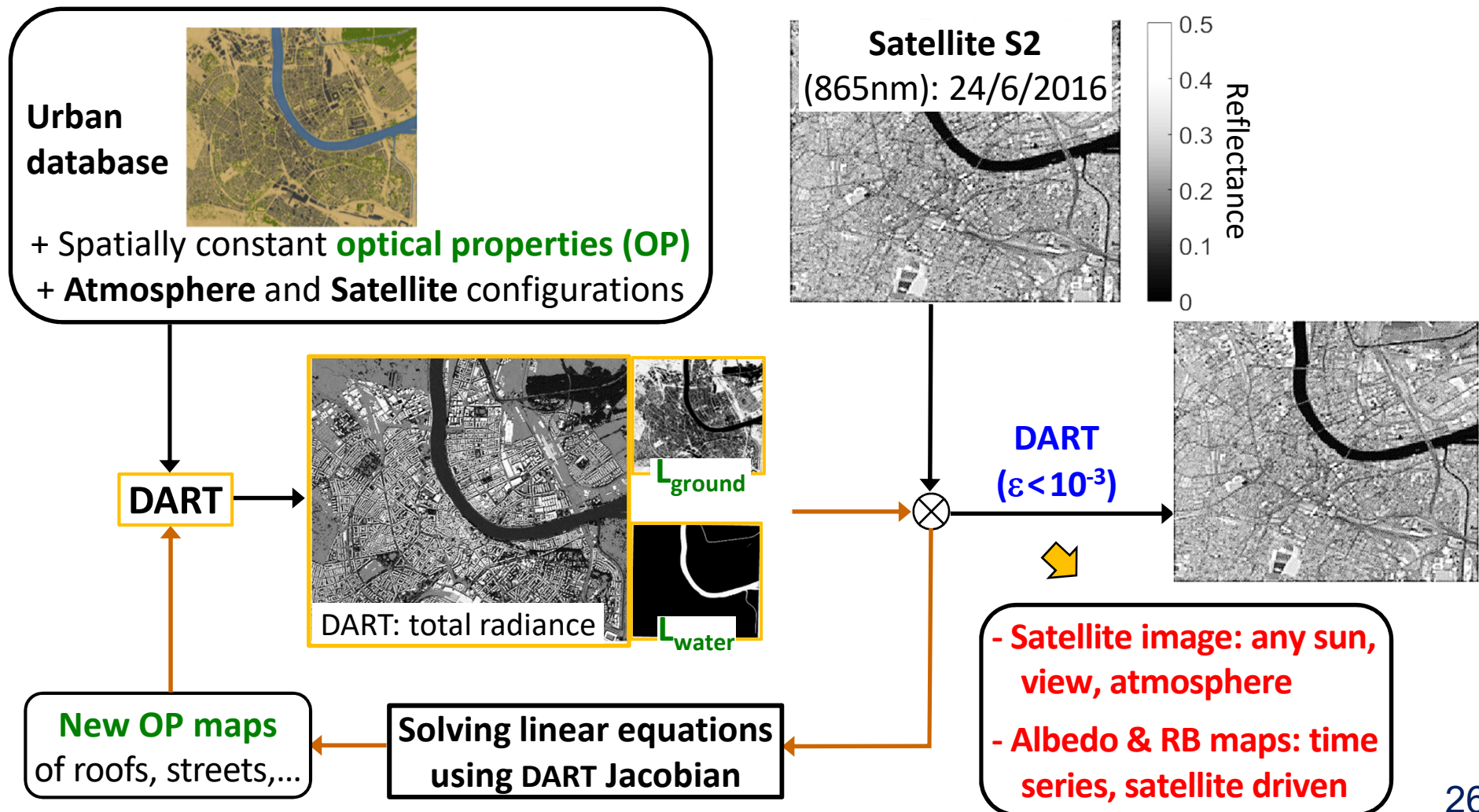


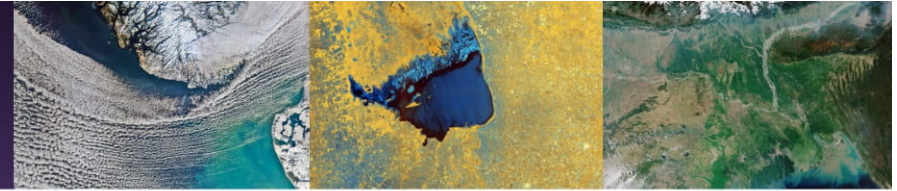
Jitter simulation



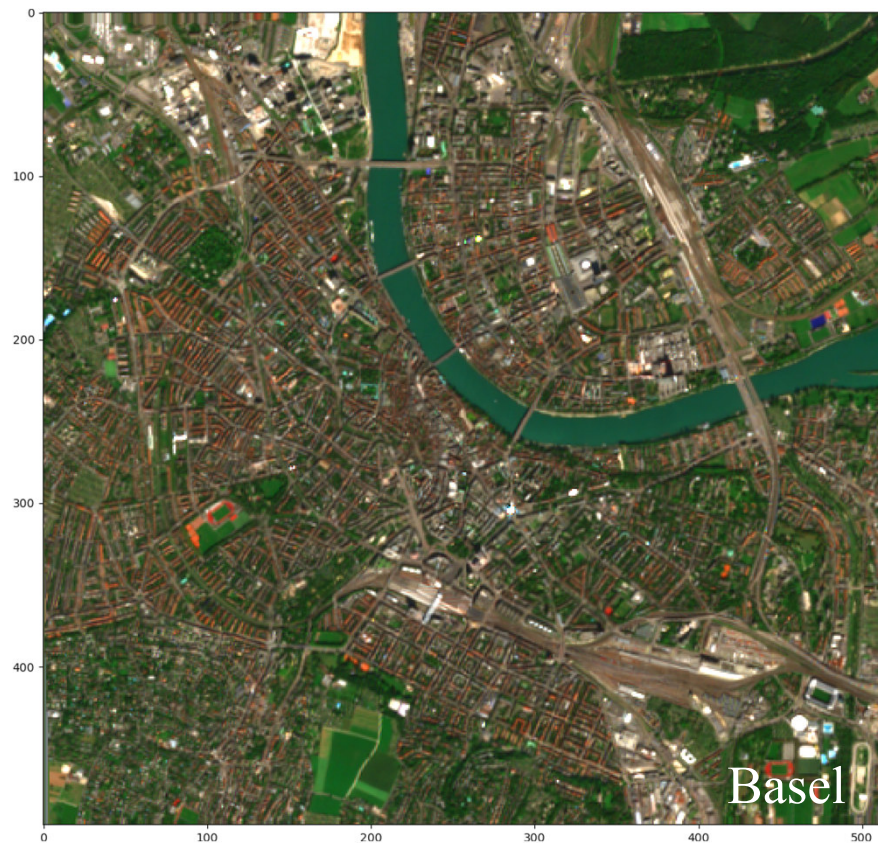


Inversion

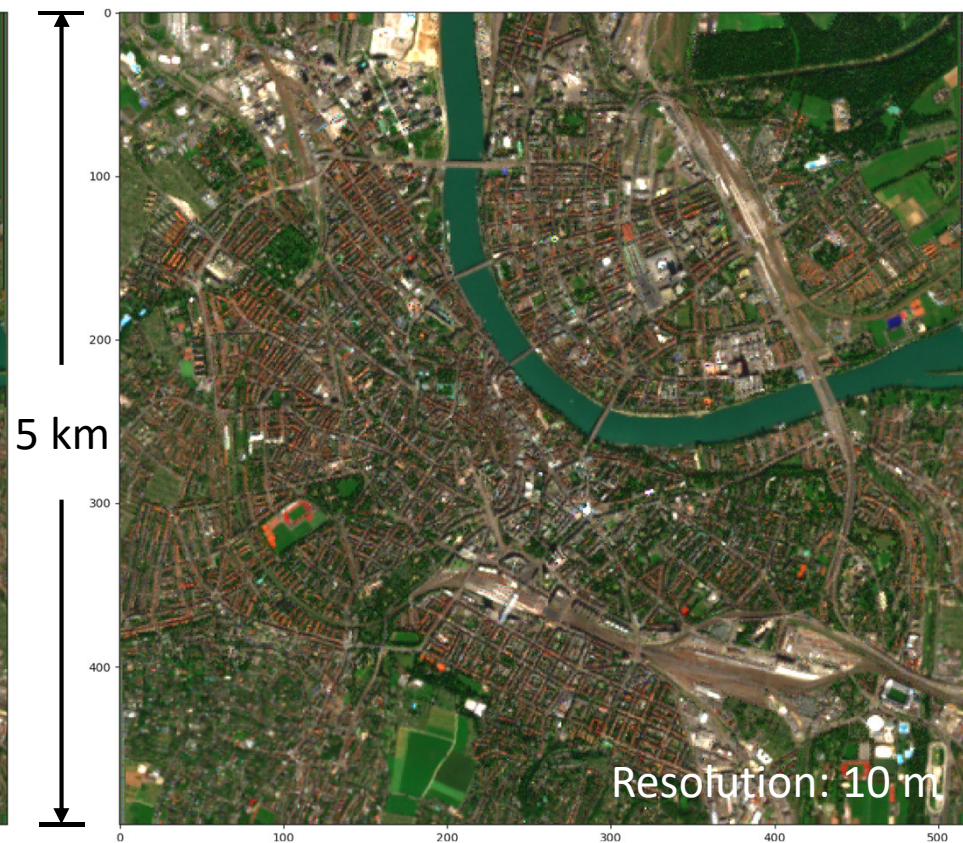




Inversion

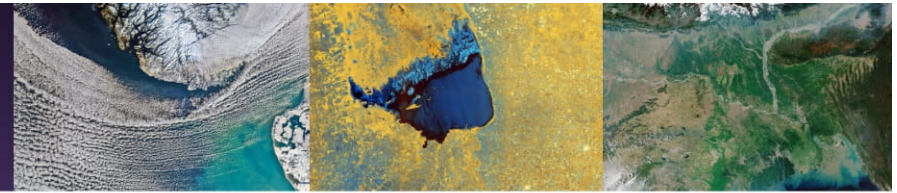


Sentinel 2 RGB (B2, B3, B4) - VIS



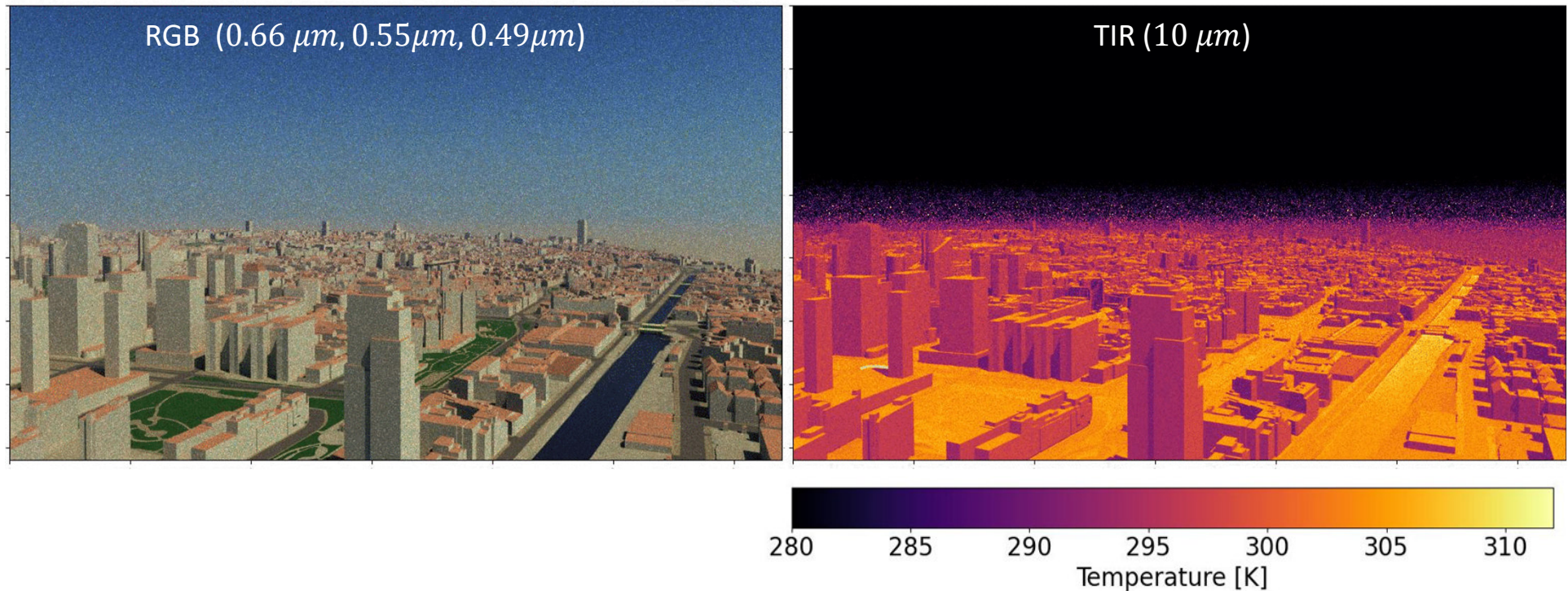
DART simulation with ρ/τ maps

(Zhen et al., 2021)

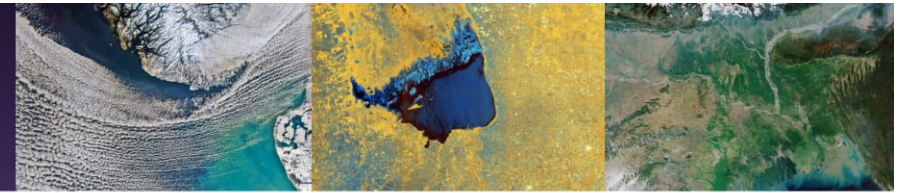


Urban heat island

Urban short wave and long wave radiation budget. [Project Suabe \(Belgium\)](#)



DART image: Brussels city. Land parameters from inversion S2/ASTER



Fire

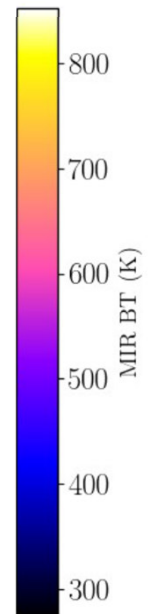
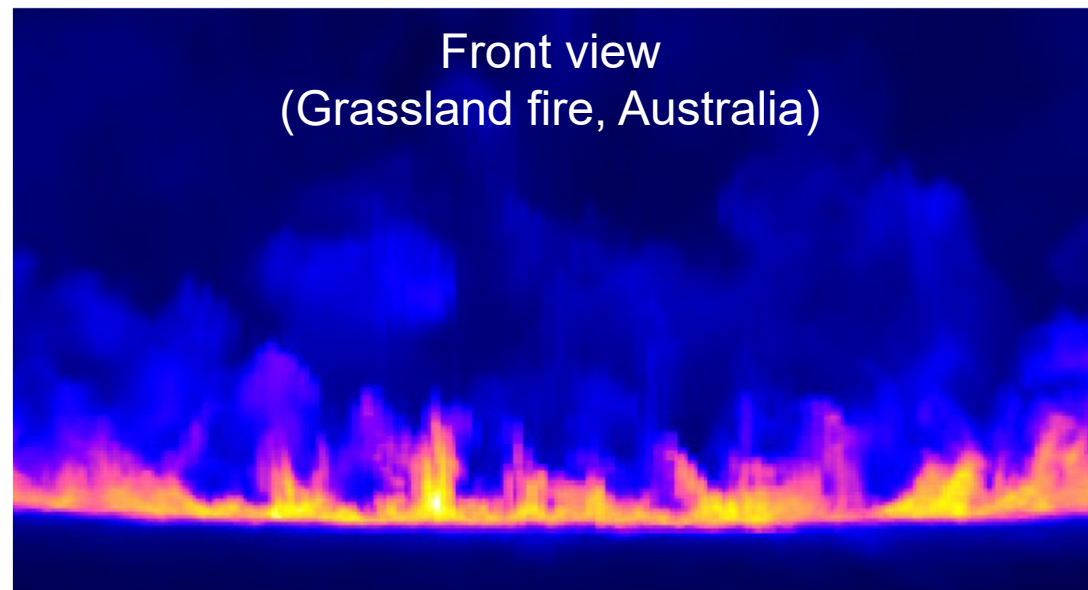
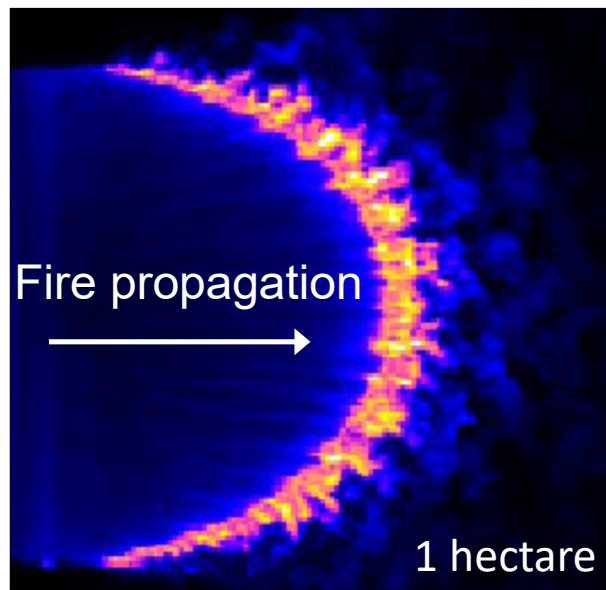
FDS (Fire Dynamics Simulator) model $\Rightarrow$ 3D temperature distribution
3D soot/gas density

DART $\Rightarrow$ Remote sensing observations (TIR camera, satellite)

$\Rightarrow$ Study the **fire radiative power** from satellite observation



UNIVERSITAT POLITÈCNICA
DE CATALUNYA
BARCELONATECH



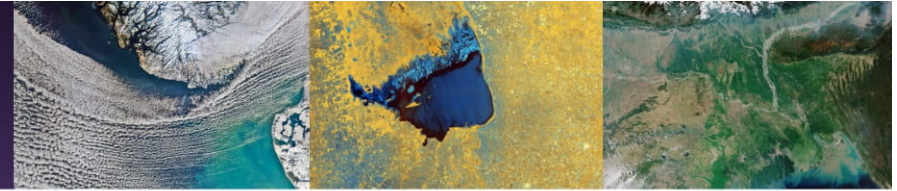
DART simulation from R. Paugam (UPC), 50s after fire ignition

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Multi-scale modelling

Forest reconstruction with TLS and QSM: 1ha, 559 trees

(Calders, et al. 2018)

(Liu et al. 2022)

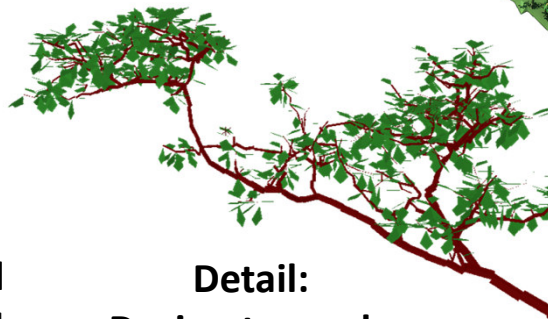
Q-FORESTLAB



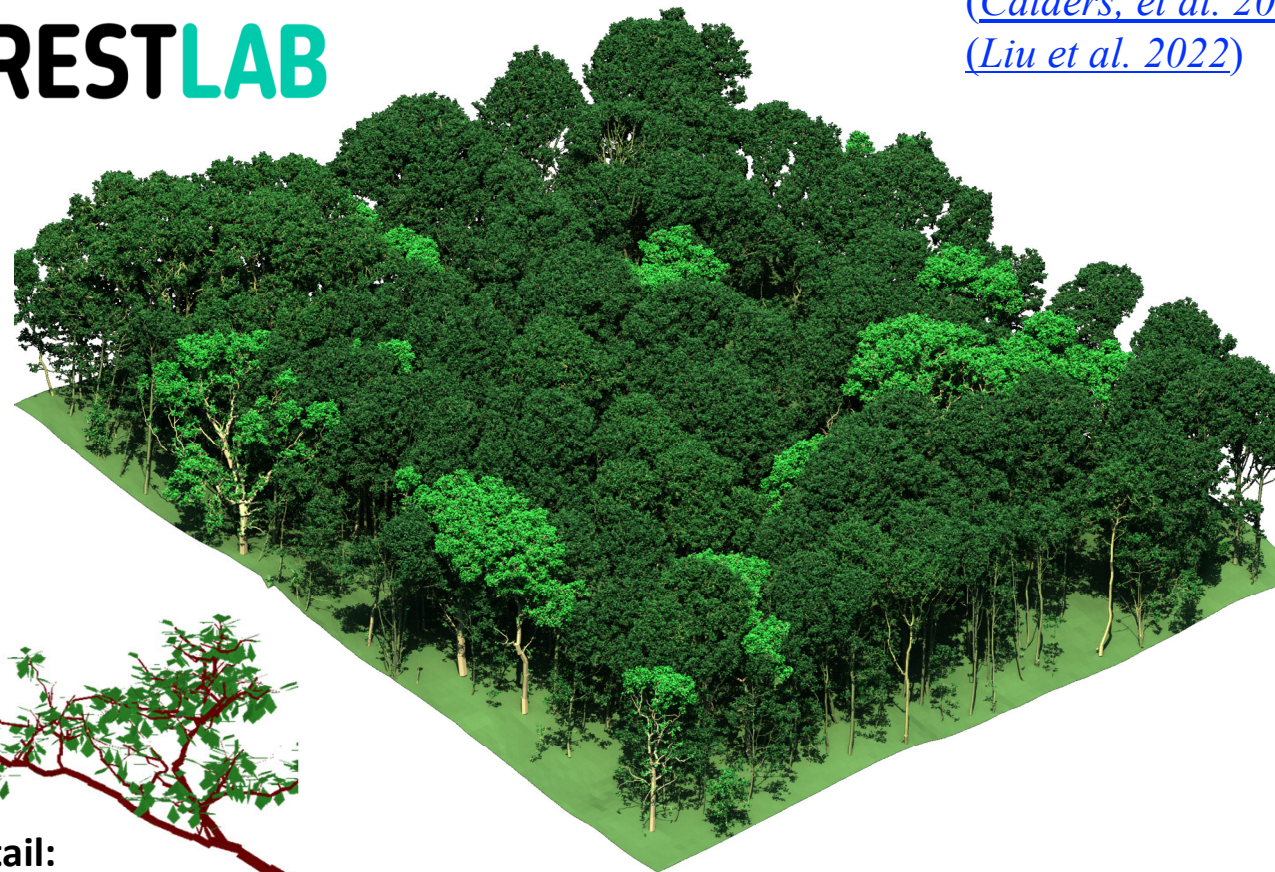
Terrestrial LiDAR
point cloud

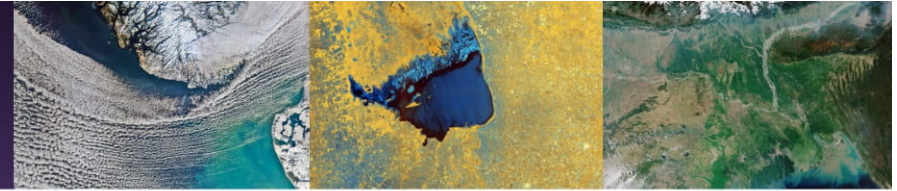


Reconstructed
3D tree model



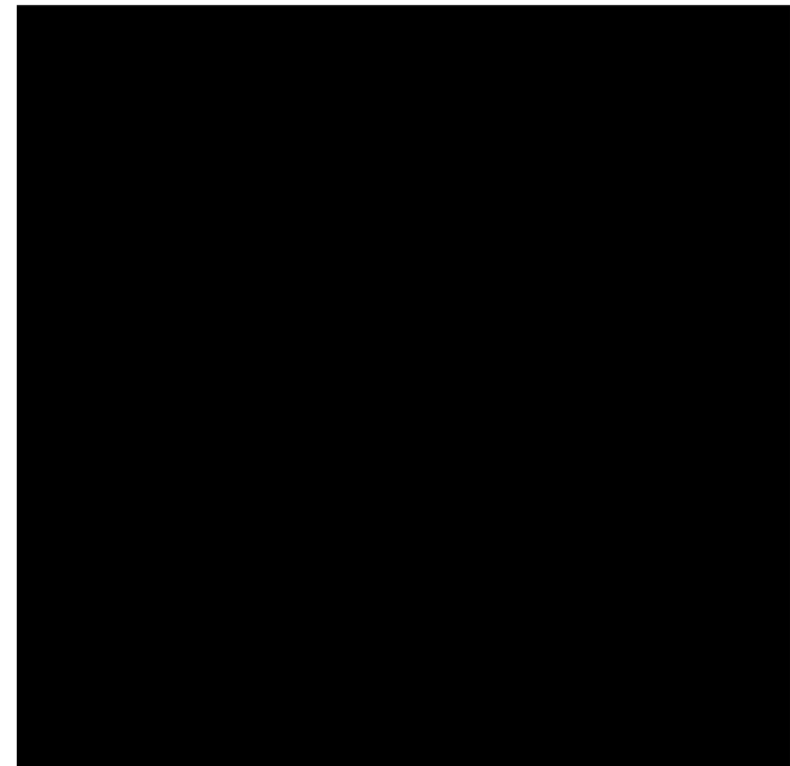
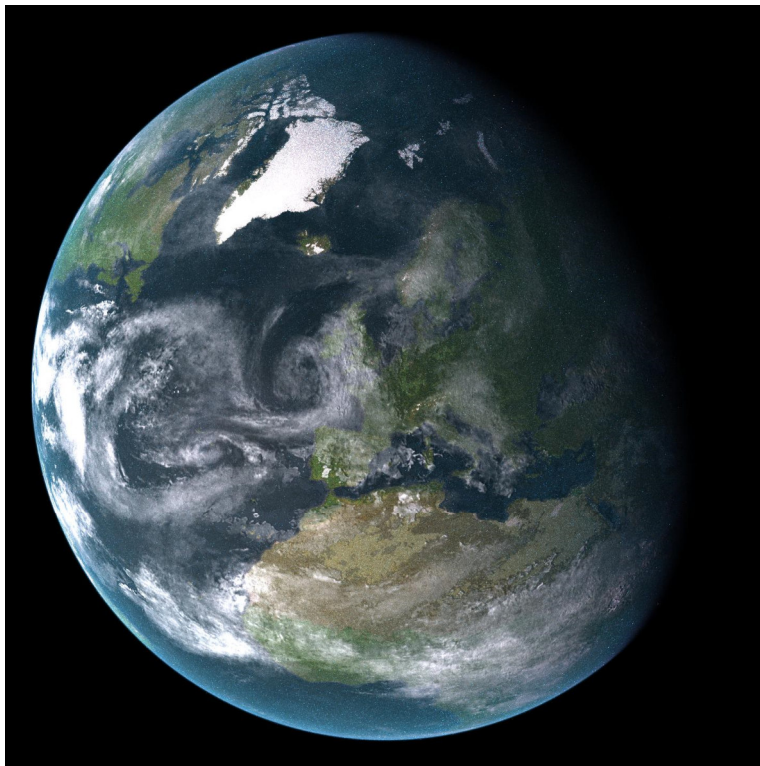
Detail:
Decimeter-scale



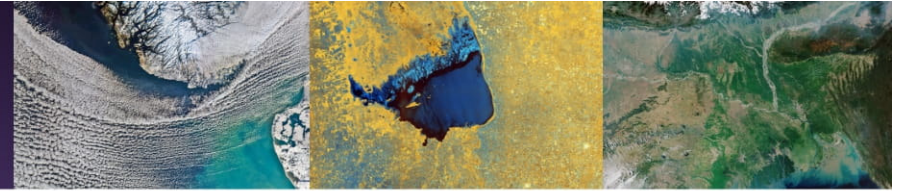


Multi-scale modelling

Data sources: Earth DEM (GEBCO 2022), Land cover (EarthEnv), Clouds cover (Copernicus EUMETSAT), Atmosphere profile (AFGL: USSTD76)



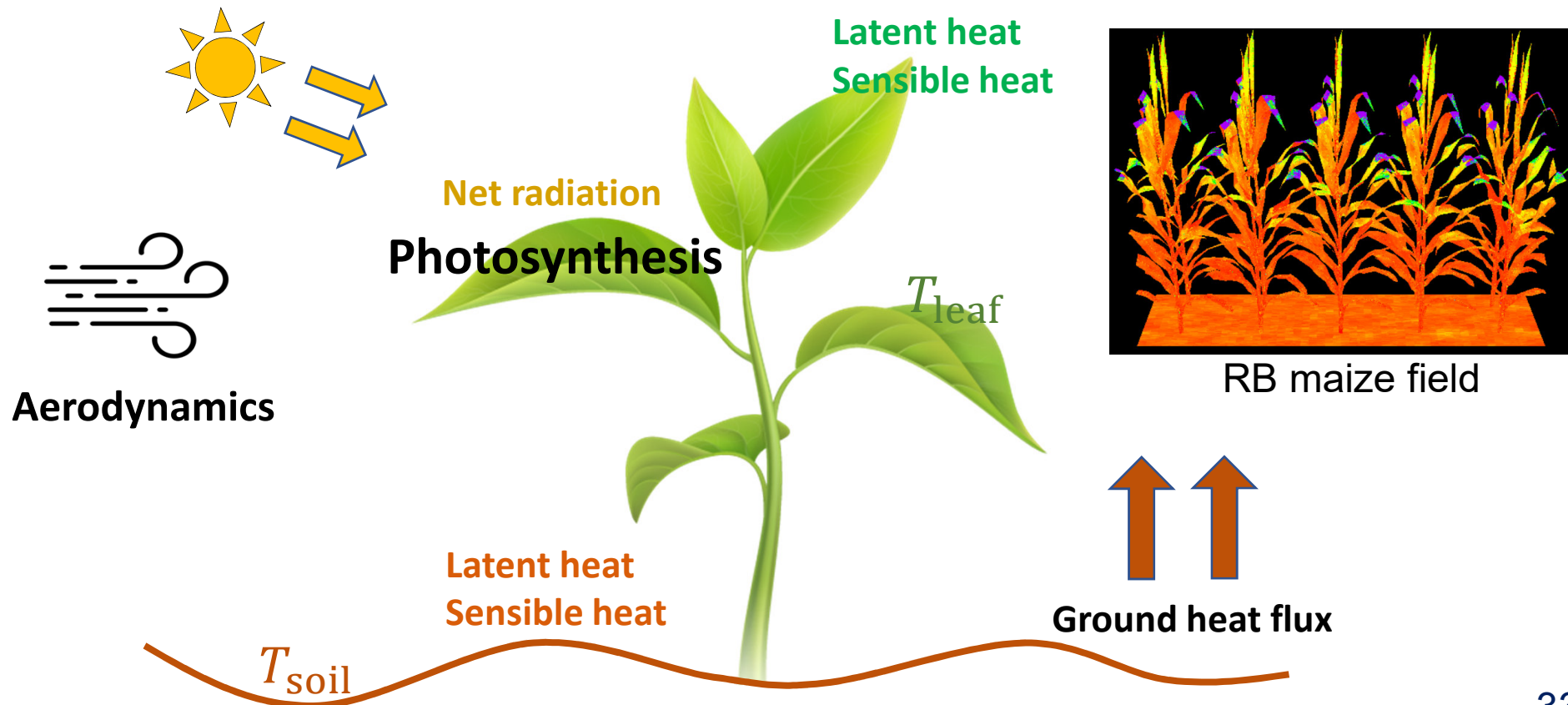
DART simulation (E. Chavanon), resolution = 10 km ($1^\circ \times 1^\circ$)

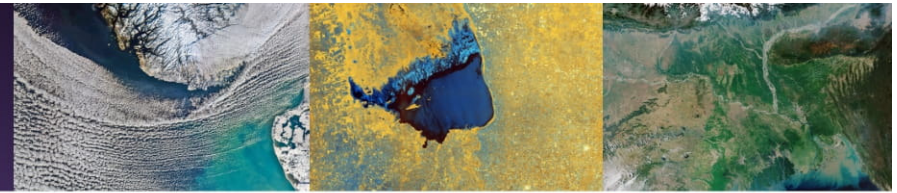


Energy balance

EB modelling: Photosynthesis, heat flux, turbulence, *etc.* + Radiative Transfer

Products: Temperature distribution, evapotranspiration, *etc.*

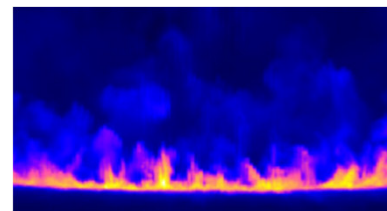




Other improvements

- (1) **Uncertainty (Jacobian):** Modelling the propagation of Uncertainties/Errors
- (2) **Georeferencing:** Import georeferenced scenes and export georeferenced images
- (3) **RT acceleration for volume:** Turbid representation of vegetation, Fire, Pollutants,...
- (4) **Measurements:** Import measurements (optical/thermal properties, 3D structure, ...)
- (5) **Surface process:** Coupling with climate/process models (SCOPE, SOLENE, ...), ...

$$\mathbf{J}_f = \begin{bmatrix} \frac{\partial f}{\partial x_1} & \dots & \frac{\partial f}{\partial x_n} \end{bmatrix}$$



THANK YOU

DART is freely available for research and education.

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Tutorial DART 17:00 – 18:00



sdrive.cnrs.fr/s/Y9bMZqGsbBKmDMf

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